



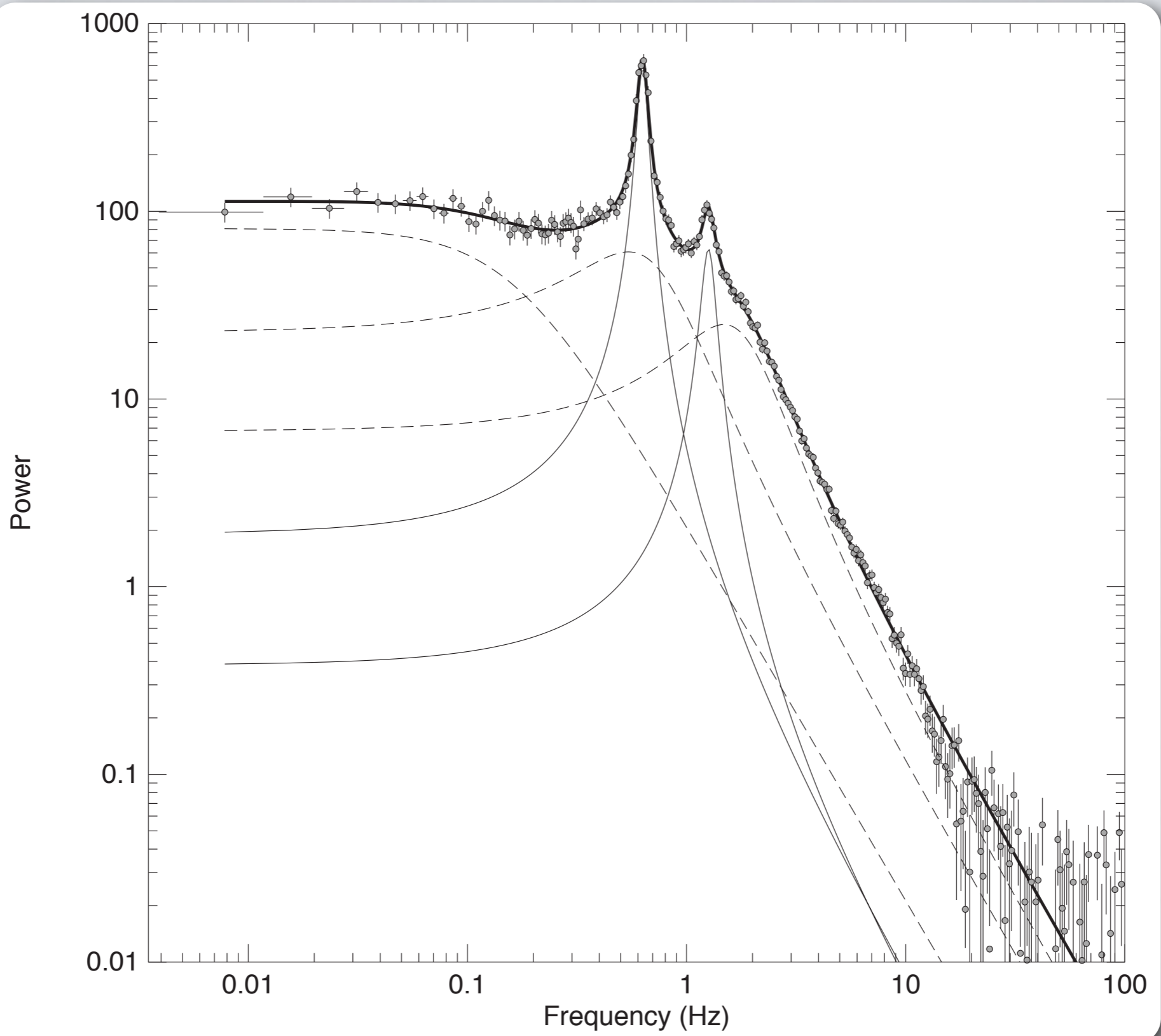
# TIME SERIES ANALYSIS

Tomaso Belloni (INAF - Osservatorio Astronomico di Brera)



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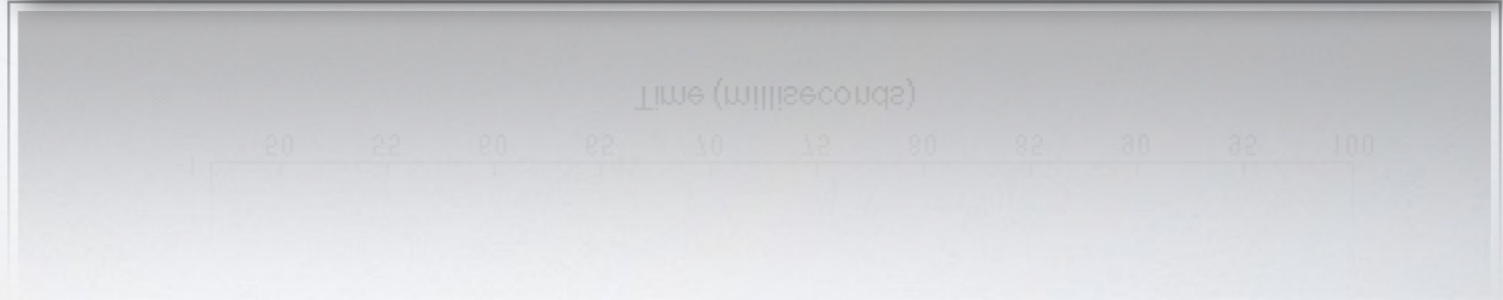
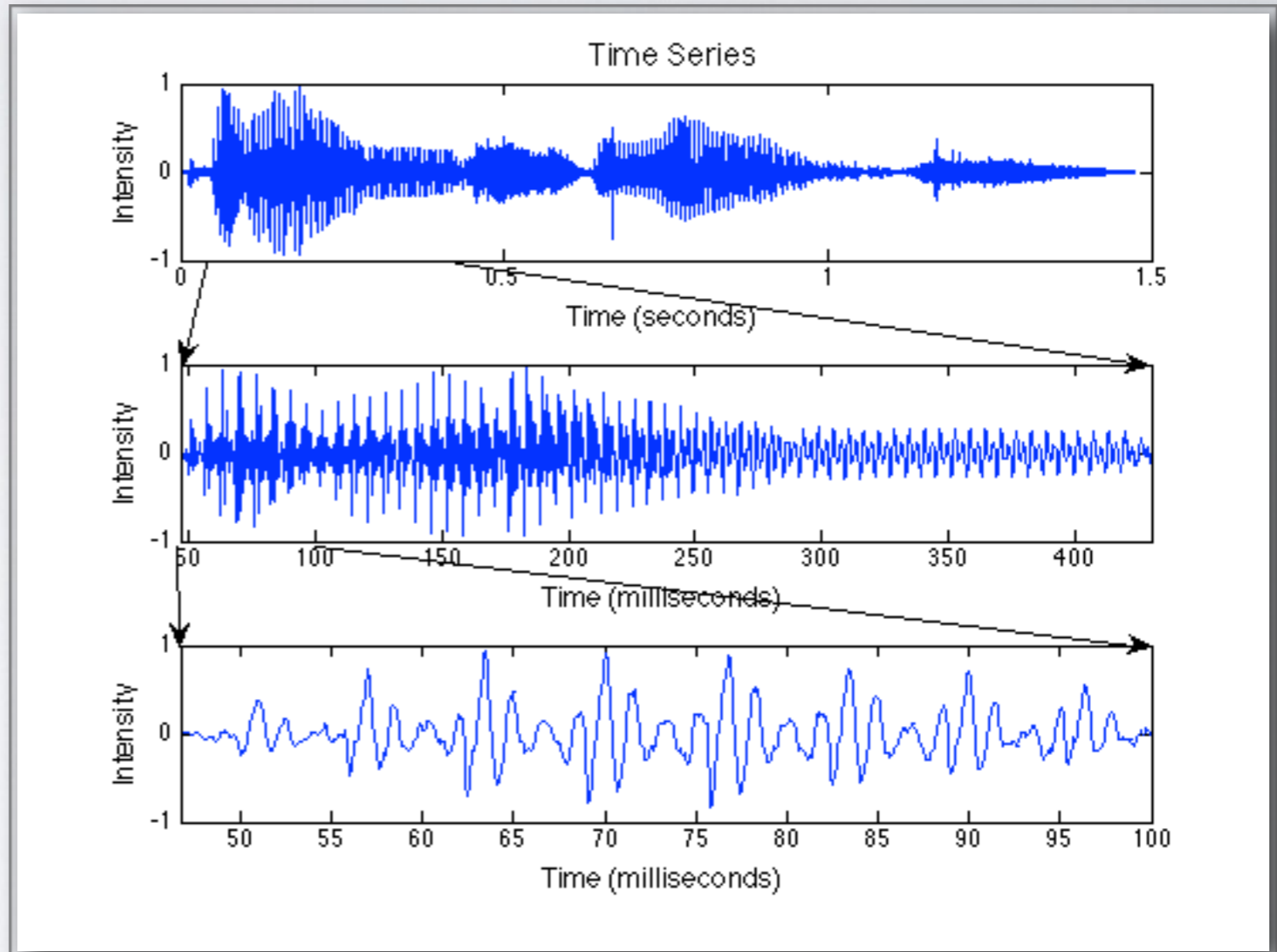
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# TIME SERIES

$h(t)$  or  $z(t)$

- I-d sequence
- Many obvious examples
- Large literature on many fields



# TIME SERIES AND FREQUENCY

Joseph Fourier (1768-1830)

- Time is important
- Different representation
- Frequency domain
- Fourier analysis
- Heat conduction → discontinuity



# FOURIER TRANSFORM

- Fourier transform equations
- $h(t)$  and  $H(f)$ : two representations of the same equation
- Linear transformation
- Decomposition on sine waves
- $\sin(2\pi f_0 t) \Leftrightarrow \delta(f-f_0)$
- Invariant to time shift

$$h(t) = \int_{-\infty}^{\infty} H(f) e^{2\pi i f t} df$$
$$H(f) = \int_{-\infty}^{\infty} h(t) e^{-2\pi i f t} dt$$

$h(t)$	$H(f)$
Real	$H(-f) = [H(f)]^*$
Even	$H(-f) = H(f)$ [even]
Odd	$H(-f) = -H(f)$ [odd]
Real & Even	$H(f)$ is real and even
Real & Odd	$H(f)$ is imaginary and odd

# OTHER BASIC PROPERTIES

- Correlation

$$\text{Corr}(g, h) = \int_{-\infty}^{\infty} g(t + \tau)h(\tau)d\tau \iff G(f)H^*(f)$$

- Autocorrelation

$$\text{Corr}(g, g) = \int_{-\infty}^{\infty} g(t + \tau)g(\tau)d\tau \iff |G(f)|^2$$

Autocorrelation is the fourier transform of the power spectrum

- Parseval's theorem

$$\int_{-\infty}^{\infty} |h(t)|^2 dt = \int_{-\infty}^{\infty} |H(f)|^2 df$$

Total power in the signal

# ONE-SIDED VS. TWO-SIDED

- Power spectral density (PSD)

$$P_h(f) \equiv |H(f)|^2 \quad -\infty < f < \infty$$

- One-sided

$$P_h(f) \equiv |H(f)|^2 + |H(-f)|^2 \quad 0 \leq f < \infty$$

- If  $h(t)$  is real

$$P_h(f) \equiv 2|H(f)|^2$$



# RECAP

$$H(f) = \int_{-\infty}^{\infty} h(t) e^{-2\pi i f t} dt$$

- Fourier transform: decomposition on a base of sinusoids
- Sum of correlation with sinusoids
- $h(t)$  extends from  $-\infty$  to  $+\infty$
- PSD over frequency gives signal power
- We have real signals...
- ... but we don't have either continuous or infinite signals

# DISCRETE FOURIER TRANSFORM

- Sampled function:  $x_k$  ( $k=1, \dots, N$ ), total length  $T$  [N numbers]

Discrete FT

$$a_j = \sum_{k=0}^{N-1} x_k e^{2\pi i j k / N} \quad (j=-N/2, \dots, N/2-1)$$

- Here times are  $t_k = kT/N$ , frequencies are  $j/T$

- Time step:  $\Delta T = T/N$

Inverse FT

- Frequency step:  $\Delta \nu = 1/T$

$$x_k = \frac{1}{N} \sum_{j=-N/2}^{N/2-1} a_j e^{-2\pi i j k / N}$$

# UNCERTAINTY PRINCIPLE (I)

- Frequency resolution:  $\Delta\nu = 1/T$
- Time resolution:  $T$  (length of the sample of  $N$  measurements)
- The longer your measurement, the higher your frequency resolution
- This is important in time-frequency analysis (non-stationary signals)
- Formal version of UP much more complex

# NO LOSS OF INFORMATION

- N numbers in input - N numbers in output (for real signals  $H(-f) = [H(f)]^*$ , but values are complex)

- Highest frequency:  $\nu_{N/2} = \frac{1}{2} \frac{N}{T}$  Nyquist frequency

- Critical sampling of a sine wave is two sample points per cycle

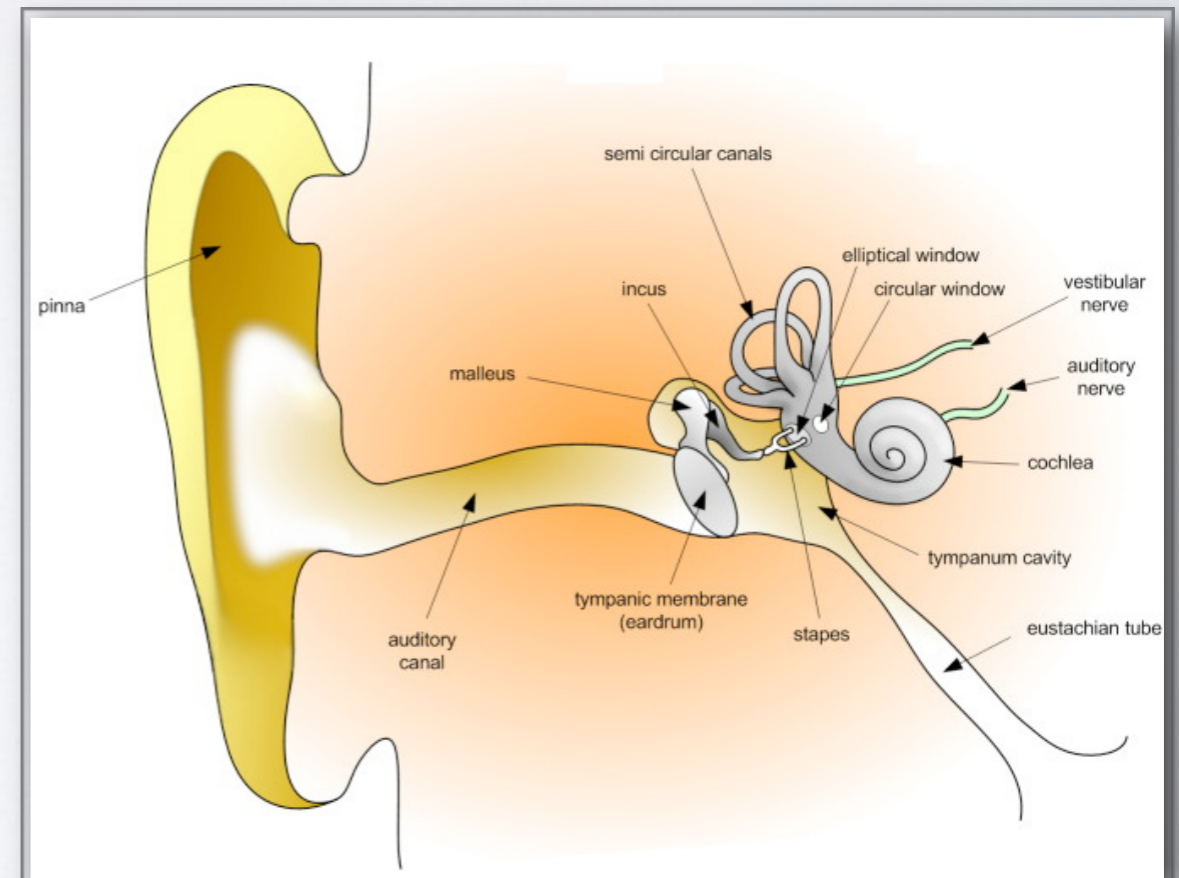
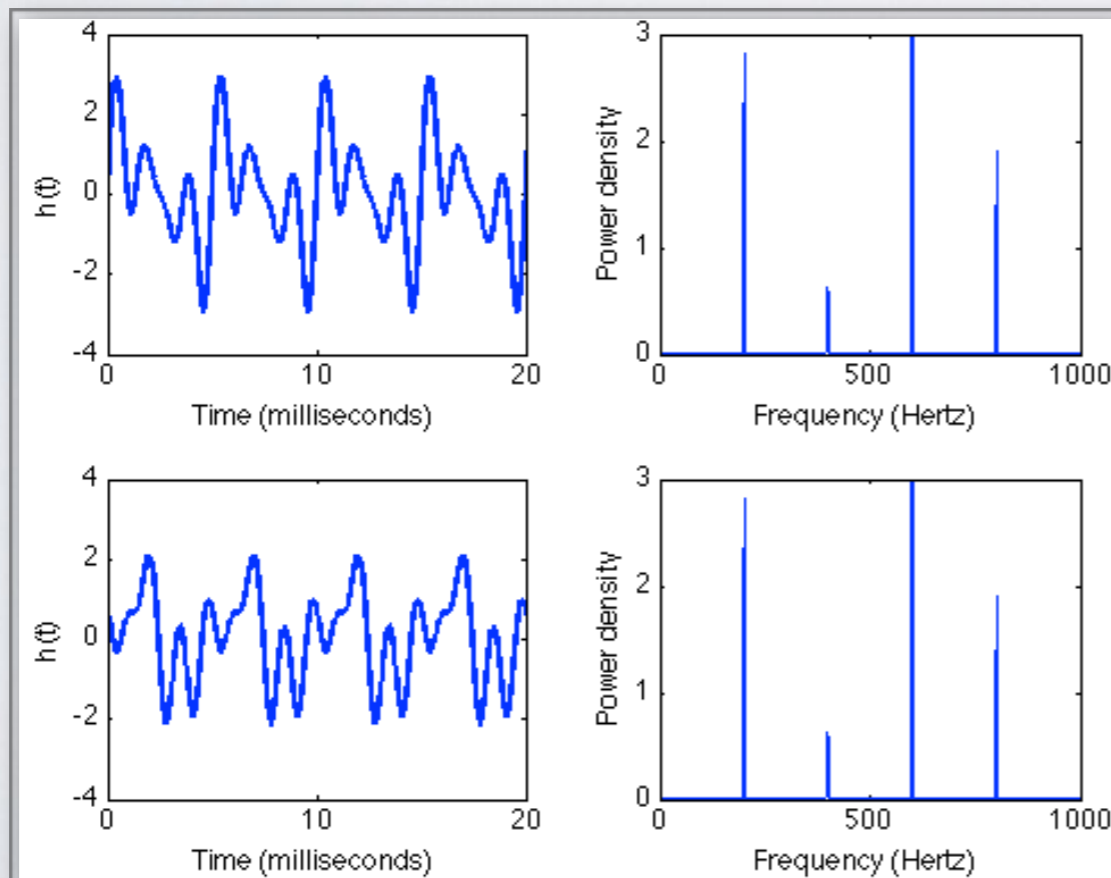
- If you sample less, you get the wrong period (wait..)

- Notice that  $H(f)$  is complex for real input

- Also: 
$$a_0 = \sum_{k=0}^{N-1} x_k e^{2\pi i 0 k / N} = \sum_k x_k \equiv N_{counts}$$

# POWER DENSITY SPECTRUM

- If we ignore the phases of the  $a_j$ 's: 
$$P = \frac{2}{N_{phot}} |a|^2 \quad (j=0, \dots, N/2)$$
- Again, analogous to hearing system

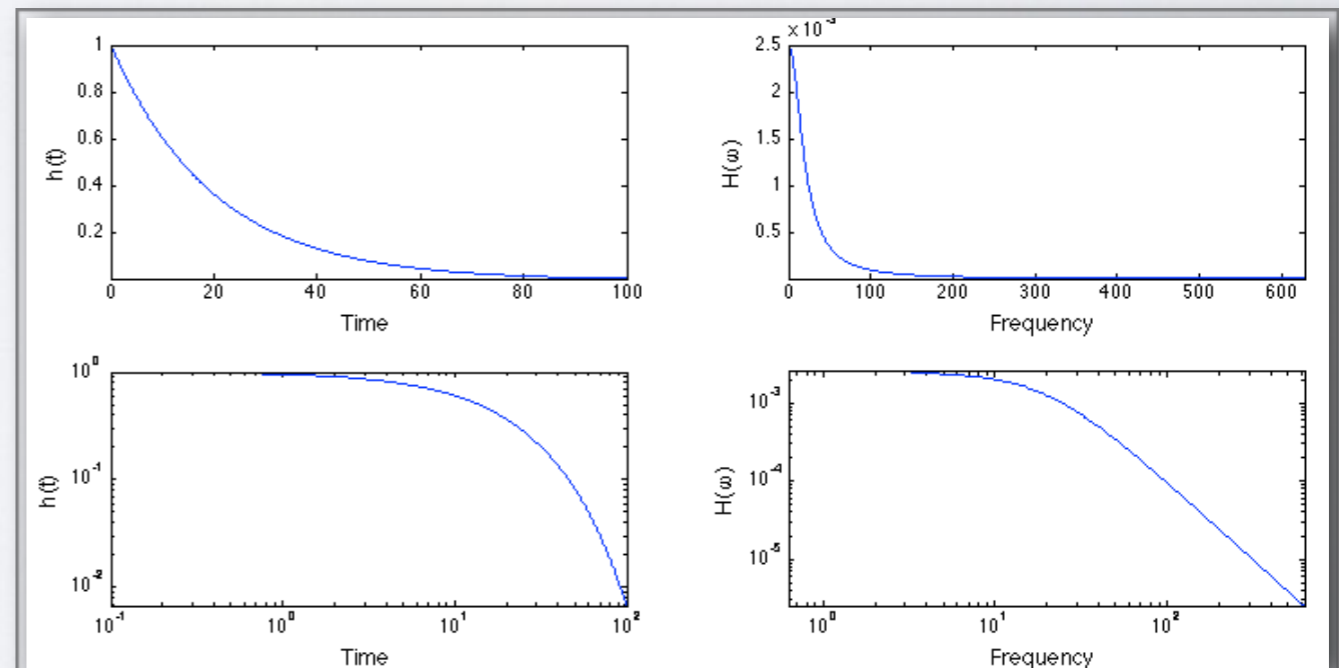


# POWER DENSITY SPECTRUM

- An example: (continuous) transform of a one-sided exponential

$$h(t) = e^{-\lambda t} \quad H(f) = \frac{1}{2\pi i f + \lambda} \equiv \frac{1}{i\omega + \lambda}$$

$$P(f) = |H(f)|^2 = \frac{1}{\omega^2 + \lambda^2}$$



# POWER DENSITY SPECTRUM

- Non-linear transformation

$$x_k = y_k + z_k \quad \longrightarrow \quad \begin{array}{l} a_j = b_j + c_j \\ |a_j|^2 = |b_j|^2 + |c_j|^2 + \text{cross terms} \end{array}$$

- If independent (random noise added),  
cross terms average out to zero

# FINITE DURATION AND SAMPLING

- How can one connect continuous and discrete FT?

$$a(\nu) = \int_{-\infty}^{\infty} h(t) e^{-2\pi i \nu t} dt \qquad a_j = \sum_{k=0}^{N-1} h_k e^{-2\pi i j k / N}$$

- Continuous time series:  $h(t) [-\infty, +\infty]$
- Discrete time series:  $h_k [k=0, \dots, N-1]$



# FINITE DURATION AND SAMPLING

- We multiply

$$h_k = h(t)w(t)i(t)$$

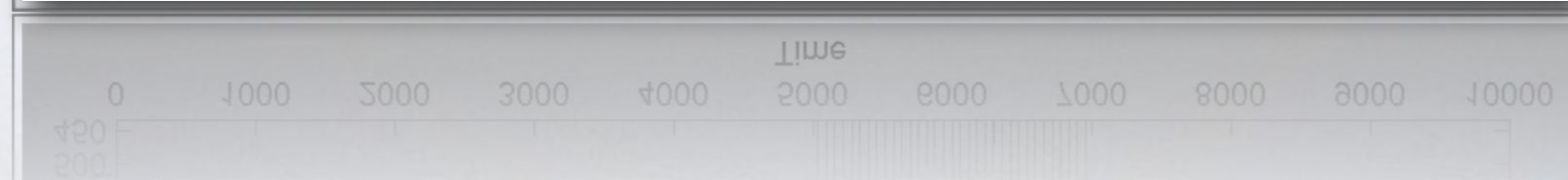
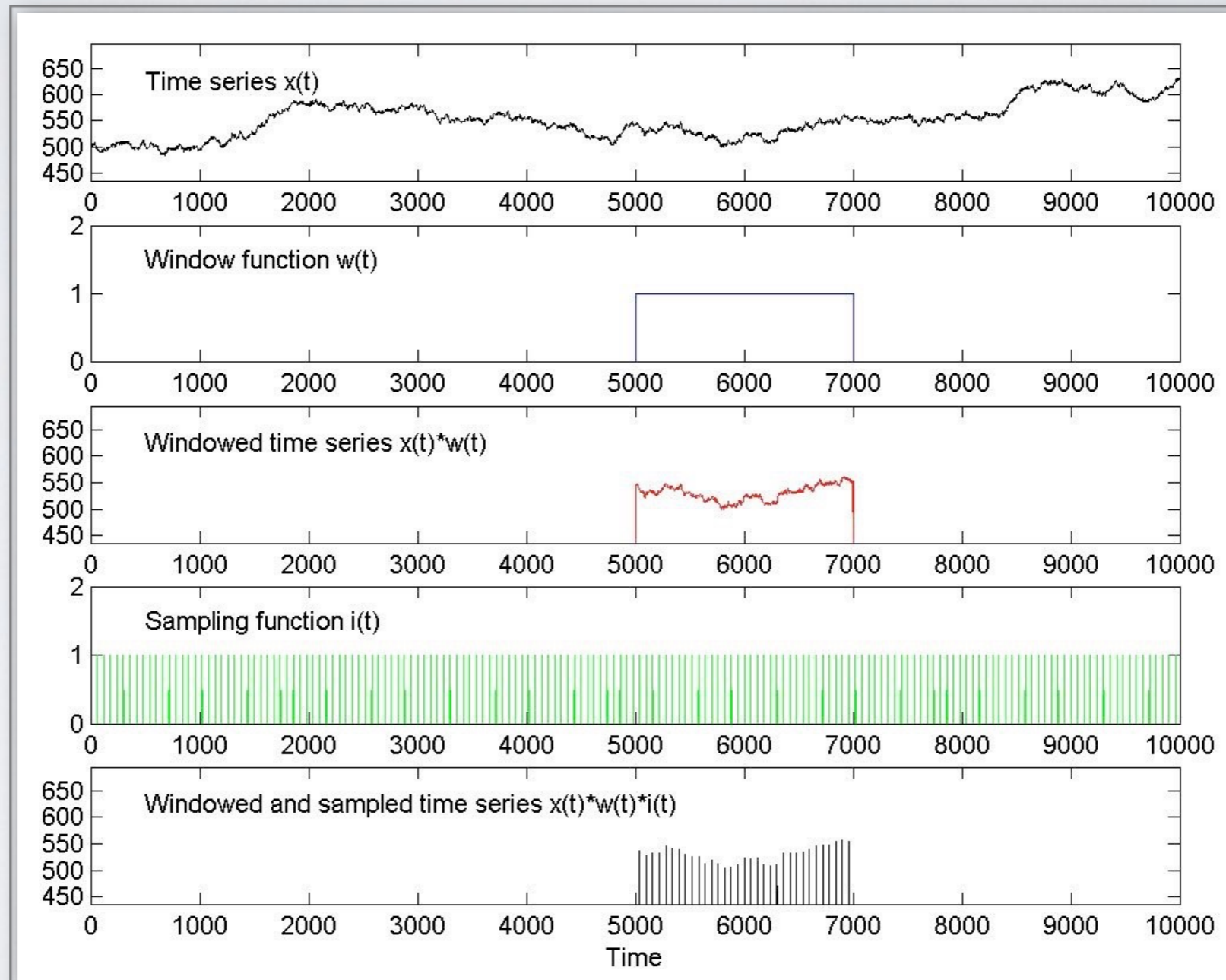
- $w(t)$ : window function

$$w(t) = \begin{cases} 1 & 0 \leq t < T \\ 0 & \textit{otherwise} \end{cases}$$

- $i(t)$ : sampling function

$$i(t) = \sum_{k=-\infty}^{\infty} \delta\left(t - \frac{kT}{N}\right)$$

# FINITE DURATION AND SAMPLING



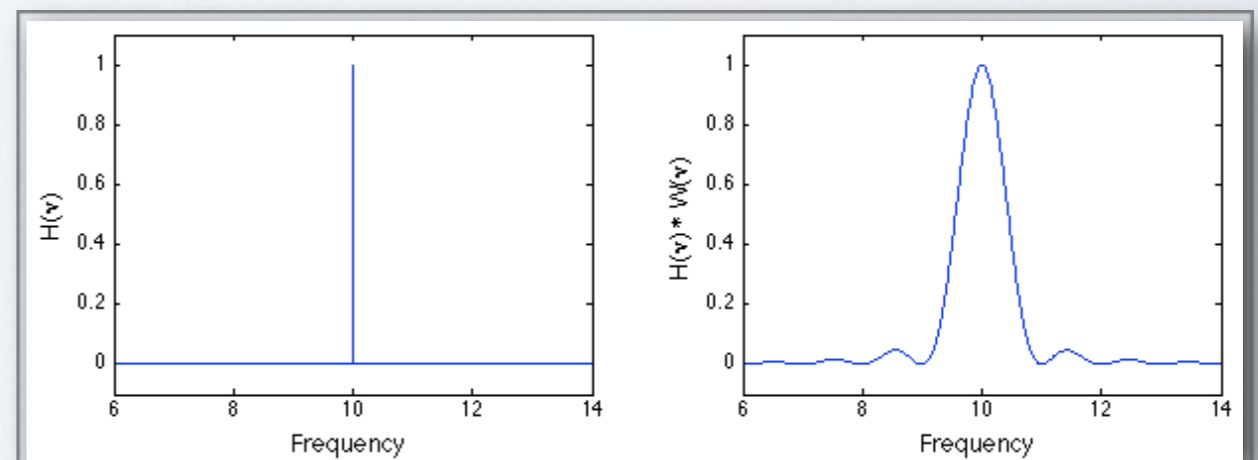
# CONVOLUTION THEOREM: WINDOWS

- The transform of the product of two functions is the convolution of the transforms

$$x(t)y(t) \iff a(\nu) * b(\nu) \equiv \int_{-\infty}^{\infty} a(\mu)b(\nu - \mu)d\mu$$

$$|W(\nu)|^2 \equiv \left| \int_{-\infty}^{\infty} w(t)e^{-2\pi\nu T} \right|^2 = \left| \frac{\sin \pi\nu T}{\pi\nu} \right|^2$$

Broadening of peaks

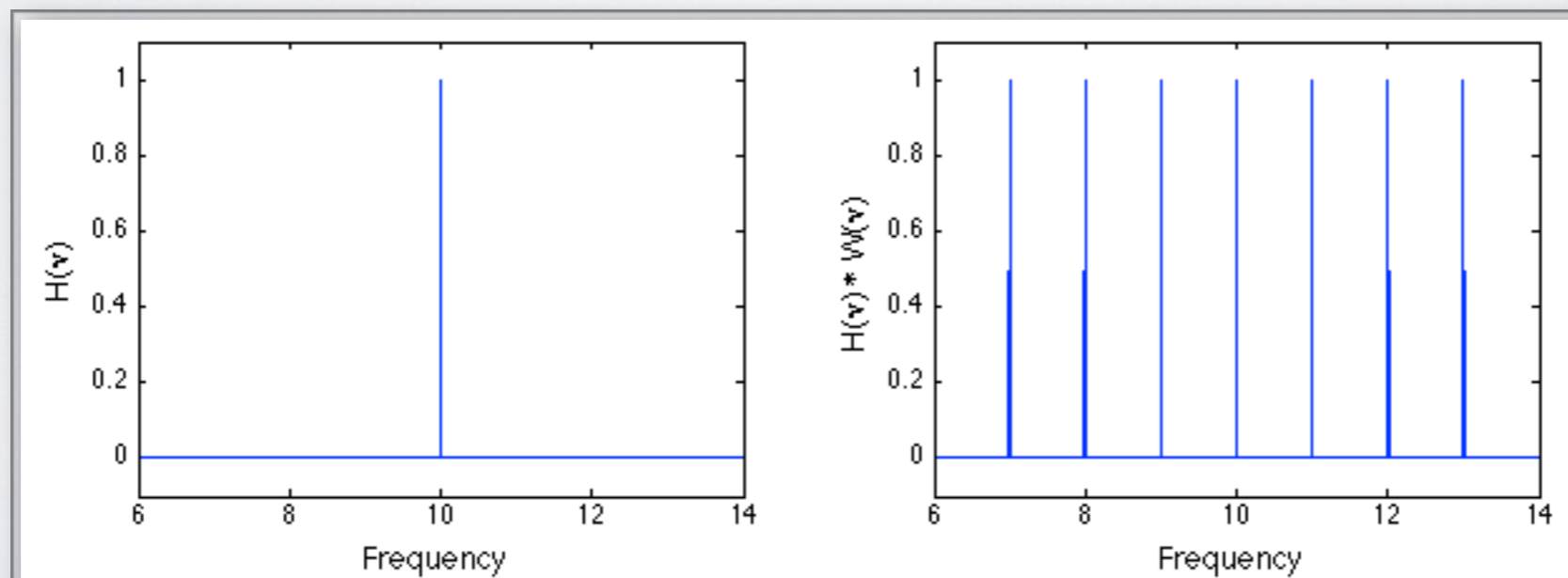


# CONVOLUTION THEOREM: SAMPLING

- The transform of the product of two functions is the convolution of the transforms

$$I(\nu) \equiv \int_{-\infty}^{\infty} i(t) e^{-2\pi\nu it} dt = \frac{N}{T} \sum_{l=-\infty}^{\infty} \delta\left(\nu - l \frac{N}{T}\right)$$

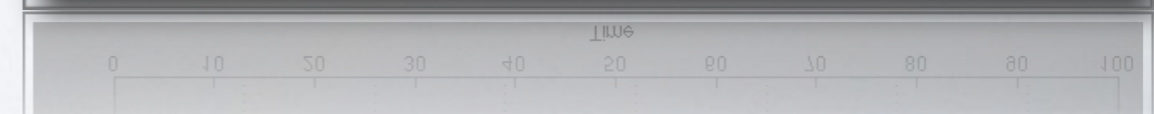
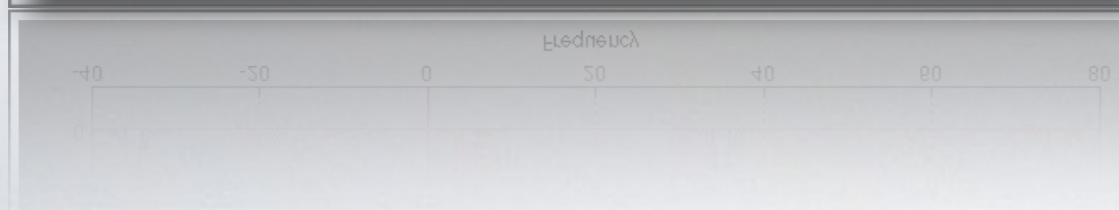
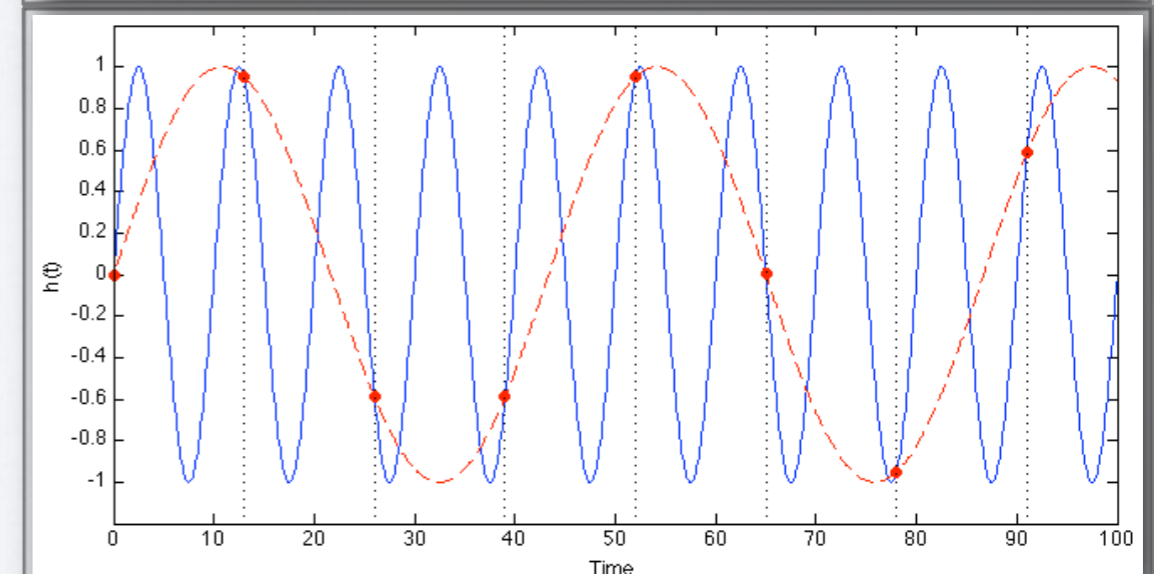
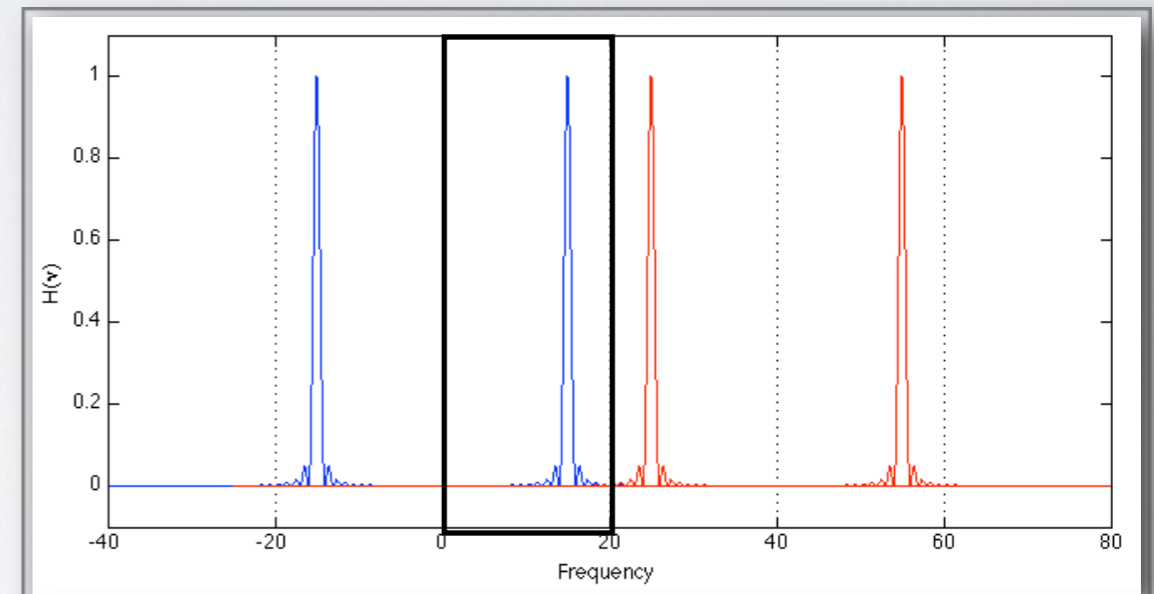
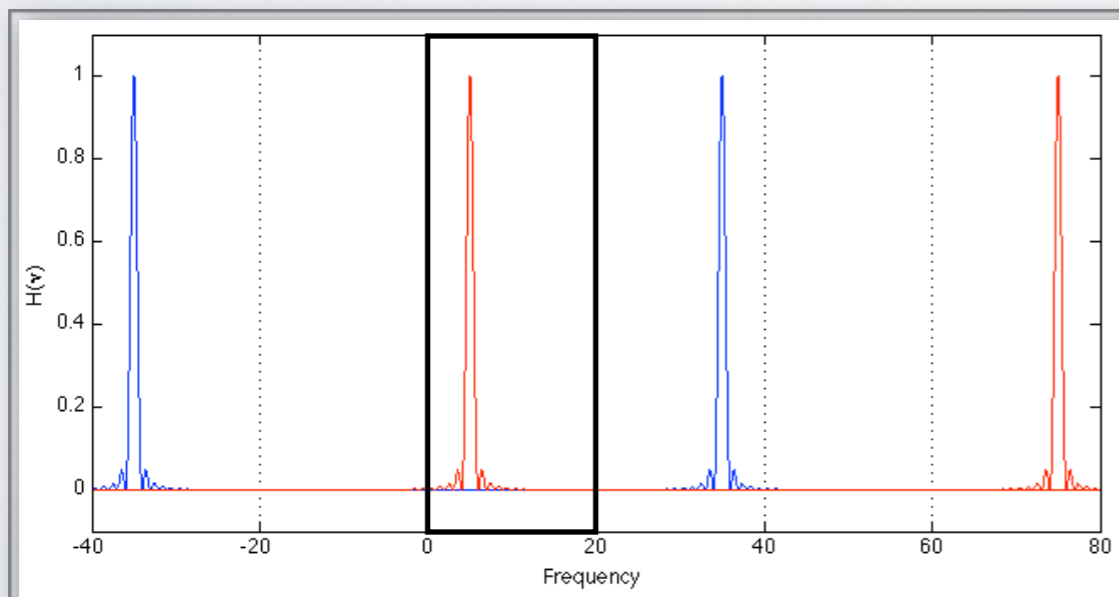
- Infinite series of  $\delta$  functions, with spacing  $N/T = 2 \nu_{Nyq}$



$> \nu_{Nyq}!$

# ALIASING

- FT is symmetric in frequency for a real signal
- Alias repeats it every  $2v_{Nyq}$
- Problem is signal above  $v_{Nyq}$



# SUMMARY OF DISCRETE FT EFFECTS

- WINDOW: broadening & sidebands
- SAMPLING: aliasing
- Aliasing not such a big problem for high-energy astronomy
- Binning, not sampling
- Suppression of high frequencies

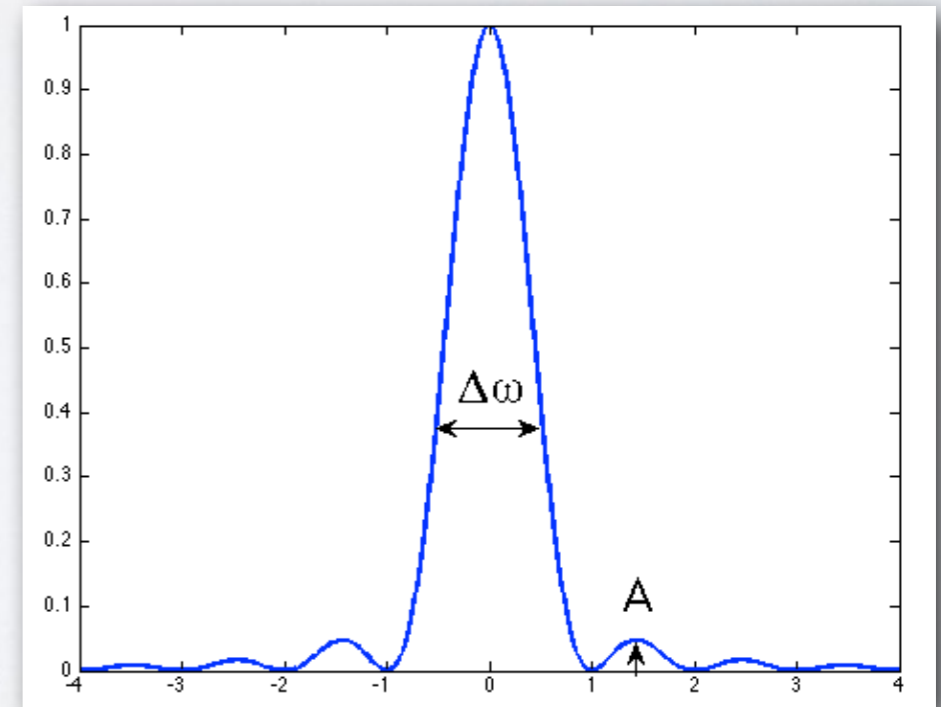
# WINDOW EFFECTS

❖ Window effect is a problem:

◆ It broadens delta peaks

◆ It flattens the slopes of noise components (sidelobes)

❖ The longer the observation, the better

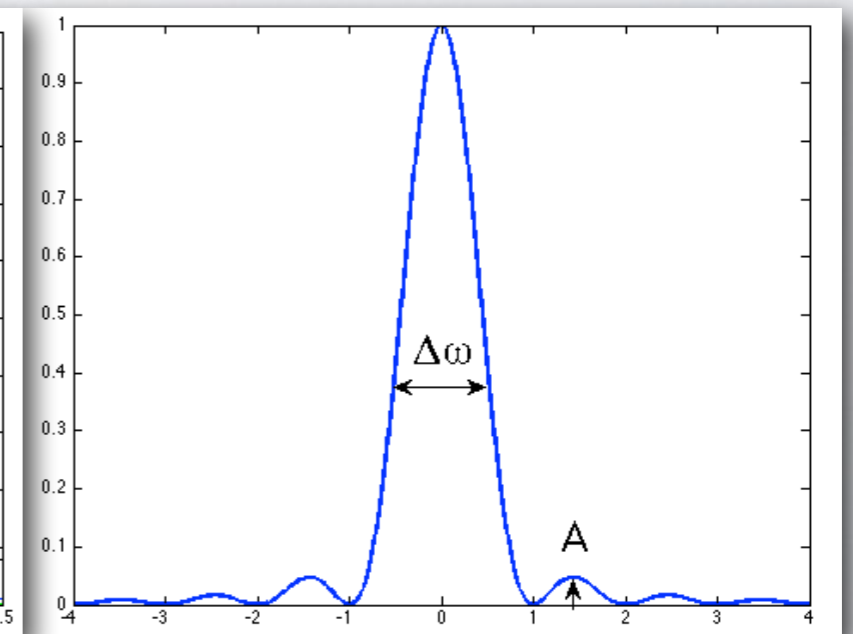
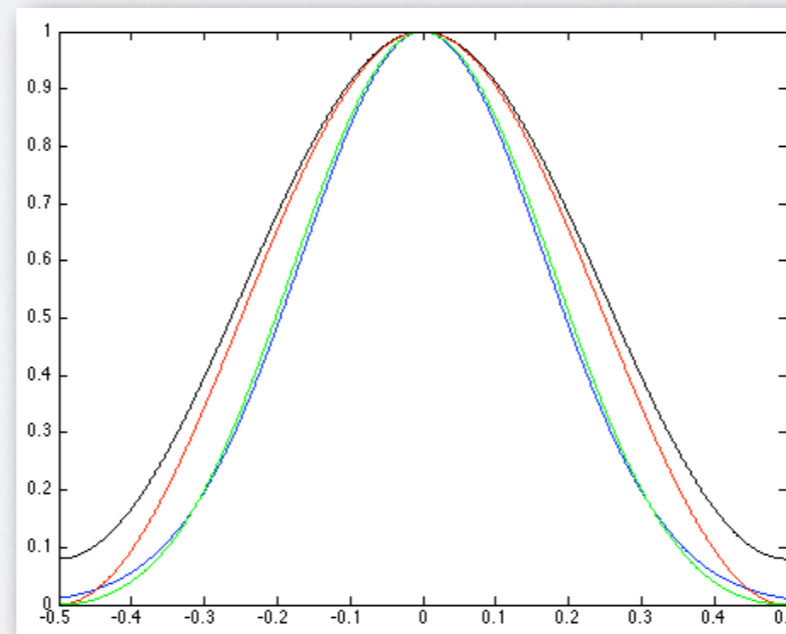


# WINDOW CARPENTRY

✿ We can use different windows

Window	$\Delta\omega$	A	p	Function
Boxcar	0.89	-13db	2	1
Hamming	1.36	-43db	2	$0.54+0.46 \cos(2\pi t)$
Gaussian	1.55	-55db	2	$\exp(-1.8t^2)$
Hanning	1.44	-32db	5	$\cos^2(\pi t)$
Blackman	1.68	-58db	5	$0.42+0.5\cos(2\pi t)+0.08 \cos(4\pi t)$

✿ We lose some signal



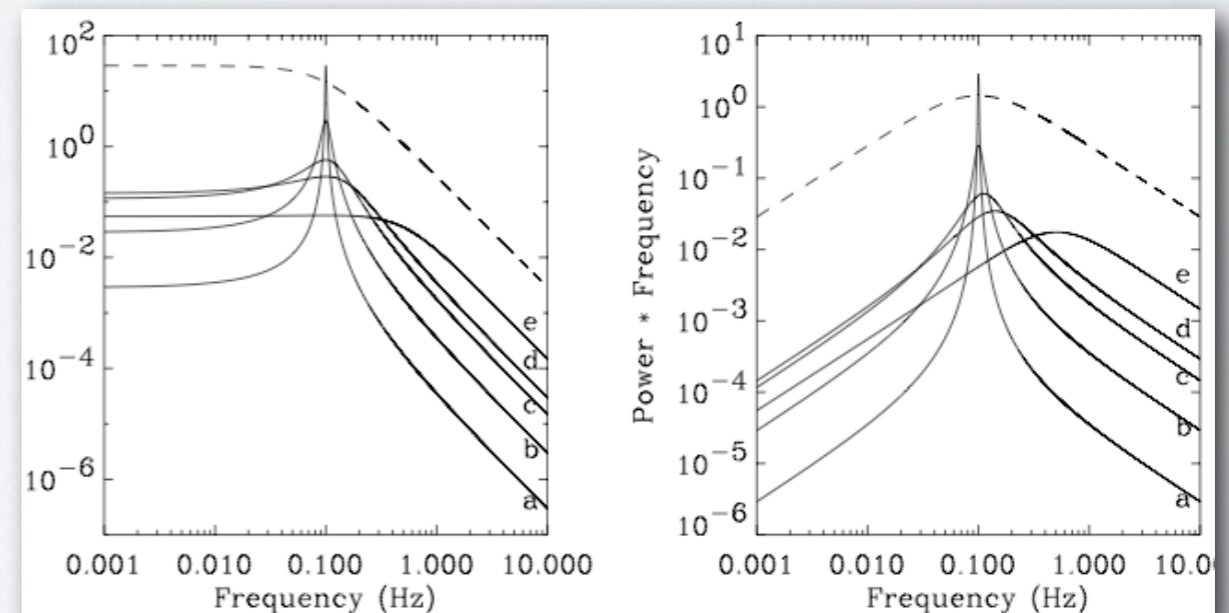


# POWER SPECTRA UNITS

- A power spectrum is in units of  $\text{Hz}^{-1}$
- It scales with the square of the intensity: variance
- If we divide by the square of the intensity, we get the fractional variance (squared rms)
- The square root of its integral is the total fractional rms
- Useful to compare amount of variability

# POWER SPECTRUM PLOTS

- Multiply the power spectrum by the frequency
- Obtain a  $\nu P_\nu$  representation
- Useful to see where the power per decade peaks
- Characteristic frequencies are peaks in  $\nu P_\nu$  (later)



Ειδικεύου (Hz)  
0.001 0.010 0.100 1.000 10.000

Ειδικεύου (Hz)  
0.001 0.010 0.100 1.000 10.0

# FAST FOURIER TRANSFORM

- What is the Fourier Transform of a single point?
- Split the series in two: odd and even points
- The FT of the series can be expressed (simply) from the FT of the two subseries
- Repeat
- Reach 1
- Reassemble

$$O(N^2) \quad O(N \log_2 N)$$

# FURTHER READING

- <http://www.brera.inaf.it/utenti/belloni/Timing/Home.html>
- M. van der Klis: Fourier Techniques in X-Ray Timing  
(technical)
- Tilman Butz: Fourier Transformations for Pedestrians  
(Springer)
- Numerical Recipes (as usual)

<http://www.brera.inaf.it/utenti/belloni/Timing/Home.html>

